

# Status of the first MiniBooNE Neutrino Oscillation Analysis

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for the MiniBooNE Collaboration

## Outline:

Intro

Janet

Event Rates

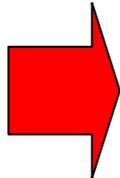
Particle Id

Bill

Backgrounds and signal

Intro: The first question....

## *Have You or Haven't You !?*



MiniBooNE has not opened the box.  
We have not yet set a date for presentation.  
We *are* in the endgame...

That's what we will show today

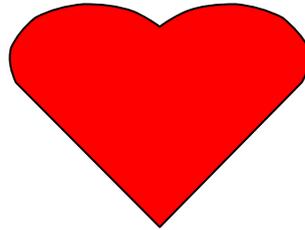
Intro: What we are asking for....

*What are we,  
the spokespersons of the neutrino run,  
asking from this PAC?*

Just a little more support and patience.

Intro: What we are asking for....

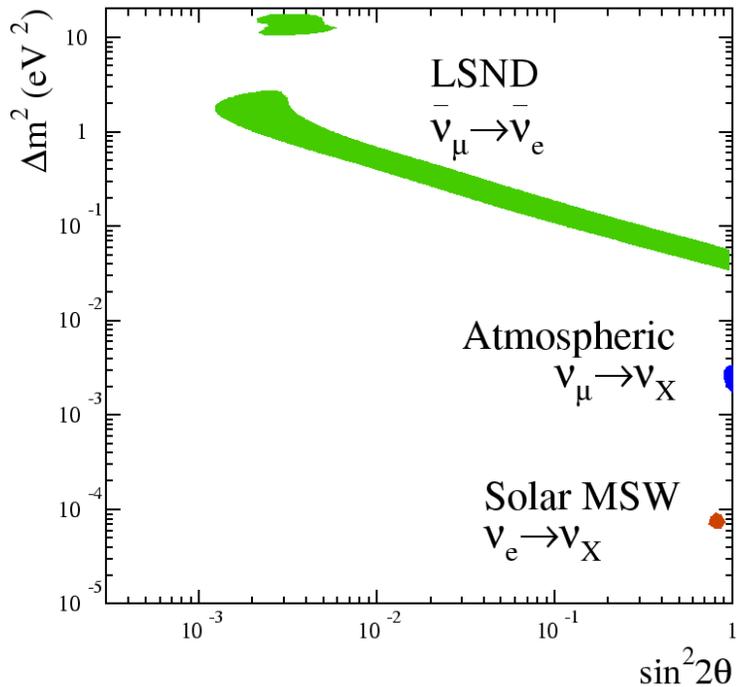
*What are we,  
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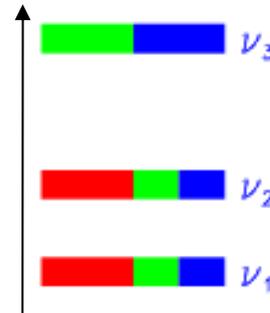
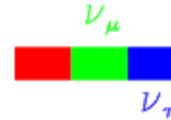
Just a little more support and patience.  
**And a shower of kudos on our collaborators  
for the great work shown here!**

# Intro: The LSND Question

**$\nu_e$  excess:**  
 **$87.9 \pm 22.4 \pm 6.0$  ( $4\sigma$ )**  
**interpreted as oscillation:**



In SM there are only 3 neutrinos



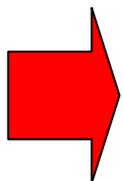
$$\Delta m_{23}^2 = m_2^2 - m_3^2$$

$$\Delta m_{12}^2 = m_1^2 - m_2^2$$

And ...  $\Delta m_{13}^2 = \Delta m_{12}^2 + \Delta m_{23}^2$

But ...  $1 \neq 0.003 + 0.00005$

LSND	atmos- pheric	solar
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Confirmation would imply a significant extension BSM

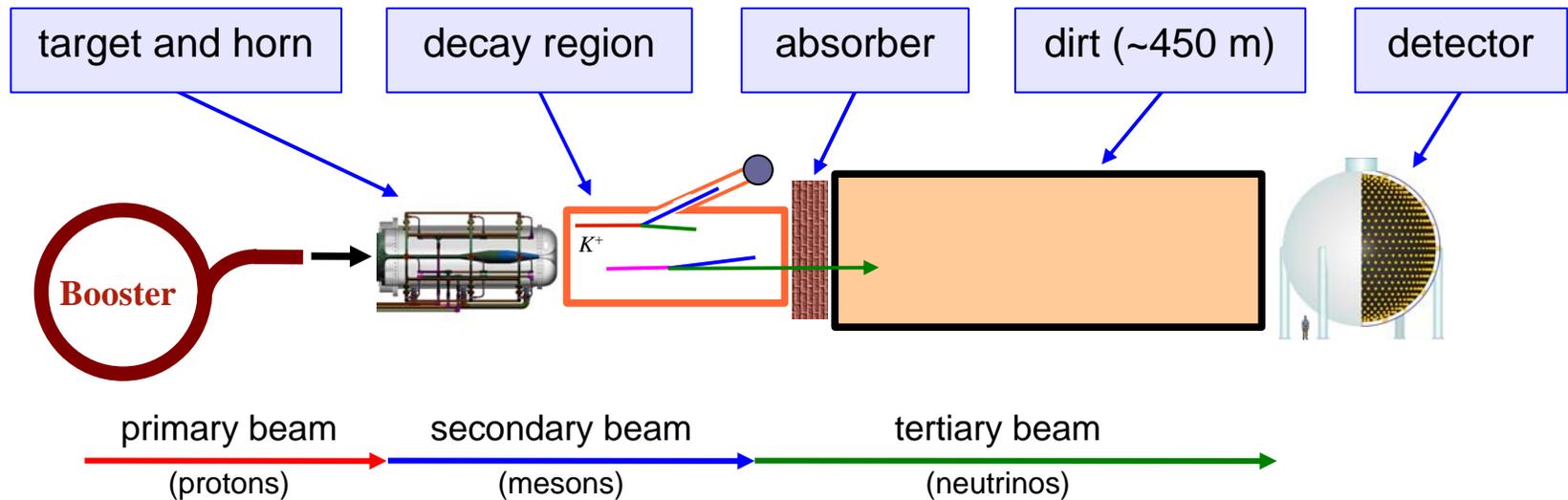
## Intro: Analysis Strategy to Test LSND

- Our goal is to test whether  
LSND is due to  $\nu_{\mu} \rightarrow \nu_e$  oscillations
- Our first result assumes no  $\nu_{\mu}$  disappearance,  
which is a good approximation given  $P_{\text{LSND}} = 0.3\%$
- We will use the  $\nu_{\mu}$  events to constrain the predicted  $\nu_e$  events  
(many systematics will cancel)
- We tie down our systematics with our own data  
External (runs in specialized set-ups)  
Internal (calibration and neutrino data)

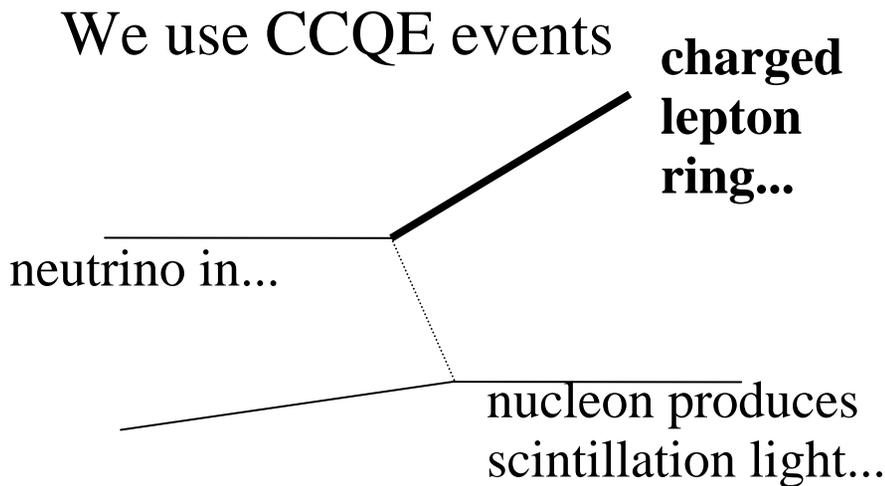
# Intro: The Design of MiniBooNE

$$P(\nu_{\mu} \rightarrow \nu_e) = \sin^2 2\theta \sin^2(1.27 \Delta m^2 L/E)$$

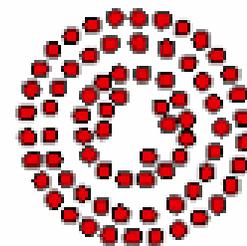
Keep  $L/E$  same while changing systematics.  
Design assumes we are looking for oscillations.



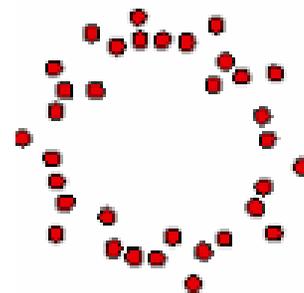
# Intro: Signal and Background Basics



**Muons**

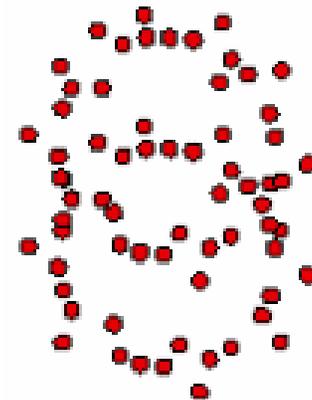


**Electrons**  
“intrinsic”  
& possibly  
from osc.



Events which look “electromagnetic”  
can be a background:

- NC  $\pi^0$  where one photon is lost
- NC radiative  $\Delta$  decay
- etc



## Event Rates: Introduction

$\nu$  Rate = flux  $\times$  cross section  $\times$  detector-related properties

Flux: From pion and kaon decays.

Use our own parameterization of outside data.

Cross section: Use the NUANCE package

With tuning to our data sets.

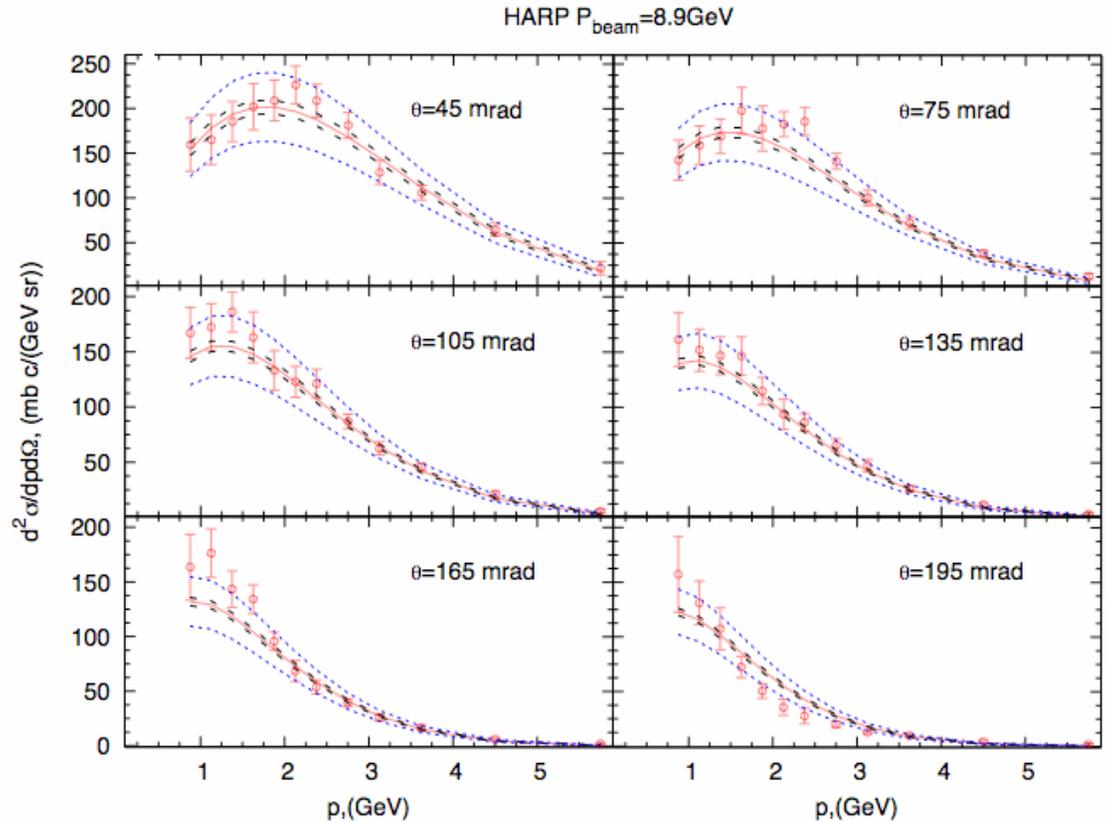
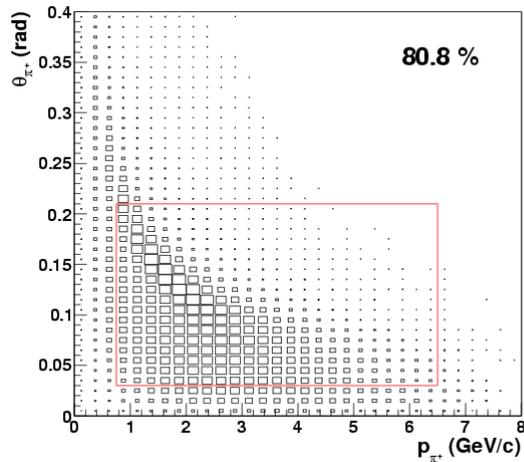
Detector: GEANT-based model.

Efficiency of reconstruction depends on  
“Optical Model” of light in the detector

*Many systematics will overlap for  $\nu_\mu$  and  $\nu_e$  in the above.*

# Event Rates: Flux Prediction from Pion Decay

Based on pBe data:  
BNL E910 @ 6,12 G  
HARP @ 8.9 GeV  
(with target replica)



Uses a “Sanford-Wang”  
Parameterization

data -- red points

total error (fit  $\oplus$  parameterization) -- blue dash

# Event Rates: Flux Prediction from Kaon Decay

$K^+$  Data from 10 - 24 GeV.

Uses a Feynman Scaling  
Parameterization developed  
by MB collaborators.

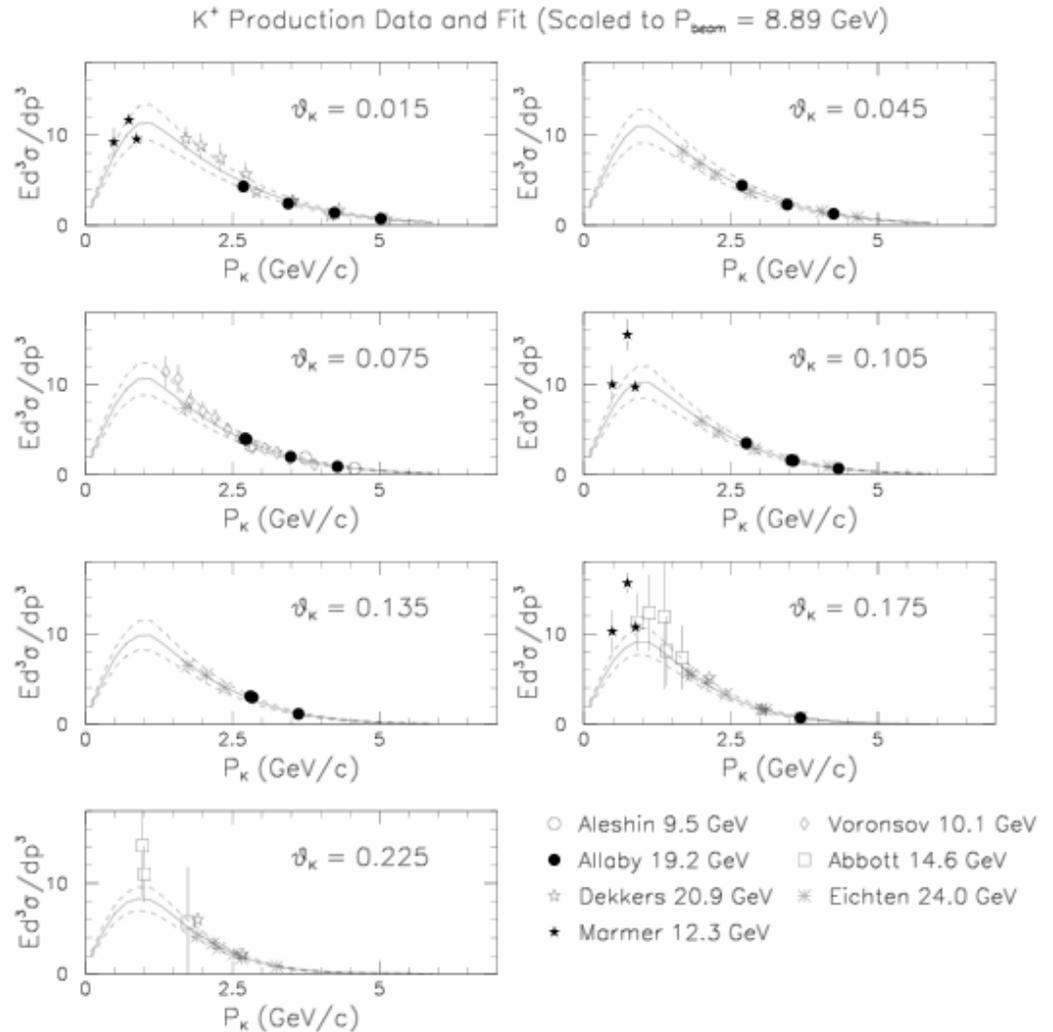
(Paper Coming Soon)

data -- points

total error

(fit  $\oplus$  parameterization) -- dash

(We also parameterize  
the  $K^0$ 's w/ a SW fit)

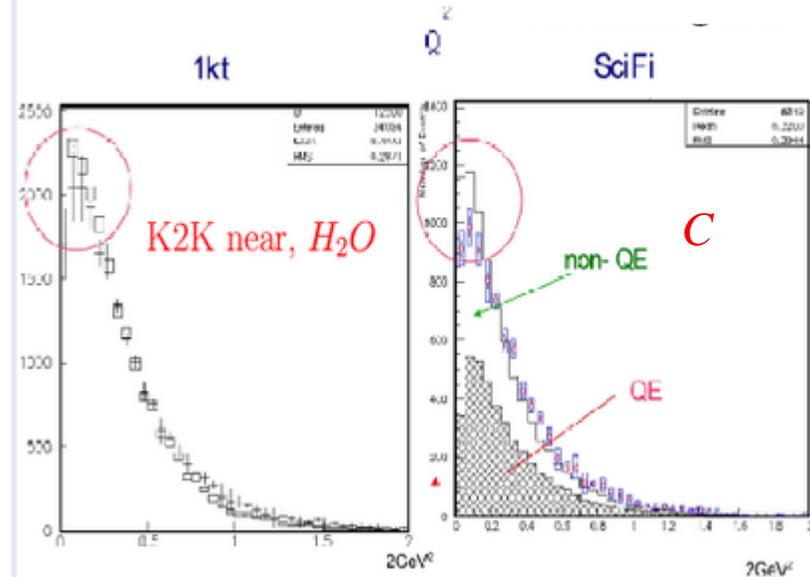
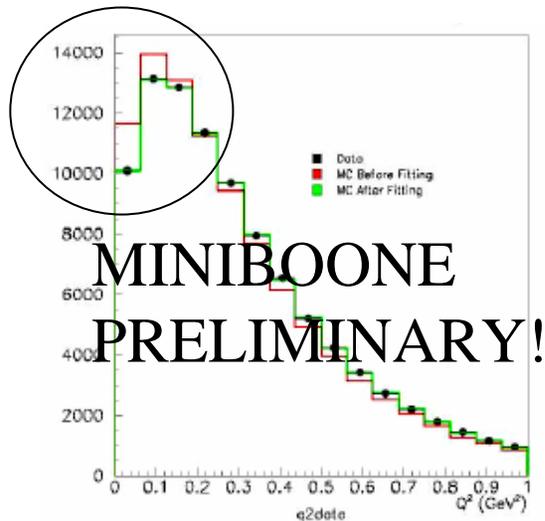


# Event Rates: CCQE Cross Section Prediction

Assumes Binding Energy and Fermi Momentum from e-scattering.

Takes the RFG as a model and fits for an effective- $M_A$  and Pauli Blocking parameter. (Paper planned)

With this, the longstanding “low  $Q^2$  discrepancy” (also seen at K2K and BNL) is addressed.



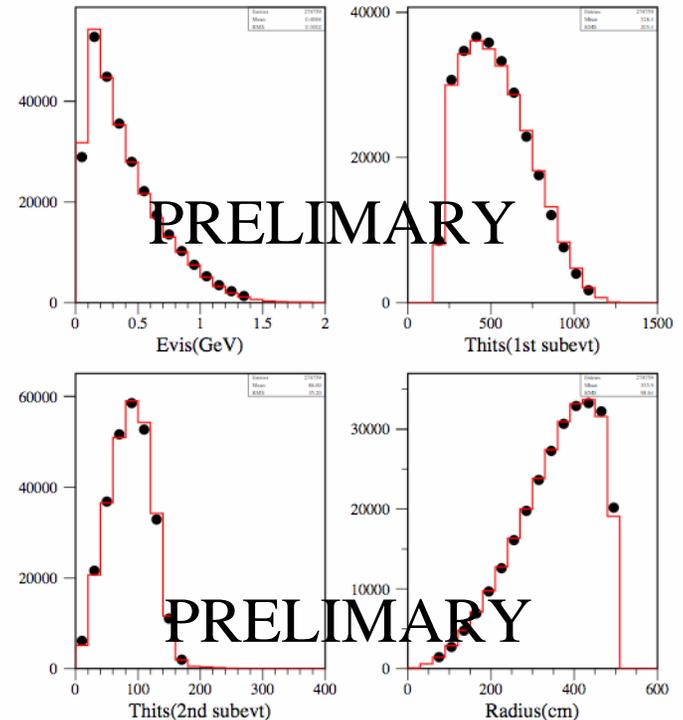
## Event Rates: Modeling the Detector

Geant simulation with a 35-parameter optical model,  
tuned using external data, Michel electrons, cosmics, NC events

Example:  
Inclusive CC events  
(tagged by Michel electron)

MC includes:

- Events in surroundings (“dirt”)
- Strobe overlay for noise, cosmics

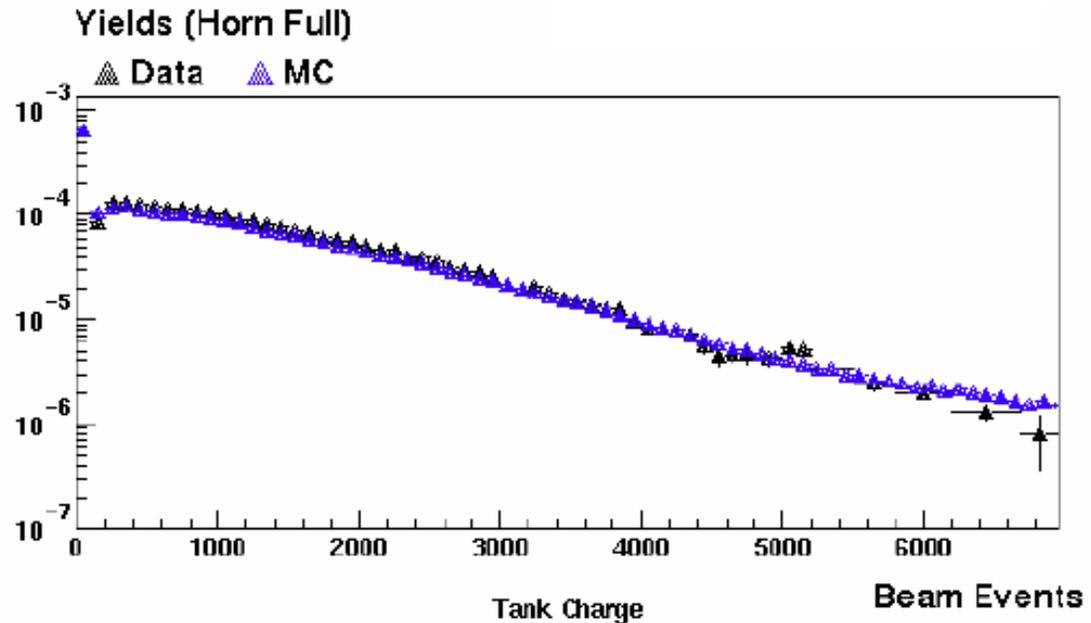


Enormous amounts of work went into tuning this MC!

## Event Rates: Comment on Overall Normalization

Data vs MC  
consistent with  
expected error

(more precise  
comment coming  
in a few weeks.)



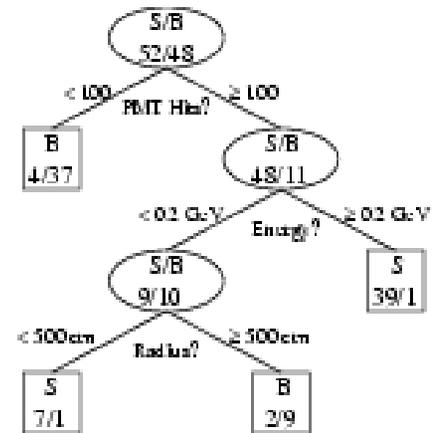
Wasn't it ~1.6 in your last PAC presentation?

Yes, improvements were all-around:

- Beam: HARP, corrected incoming beam model
- Cross Section: improved CCQE model
- Detector: improved optical model, leading to better fiducial volume agreement

# Particle Identification

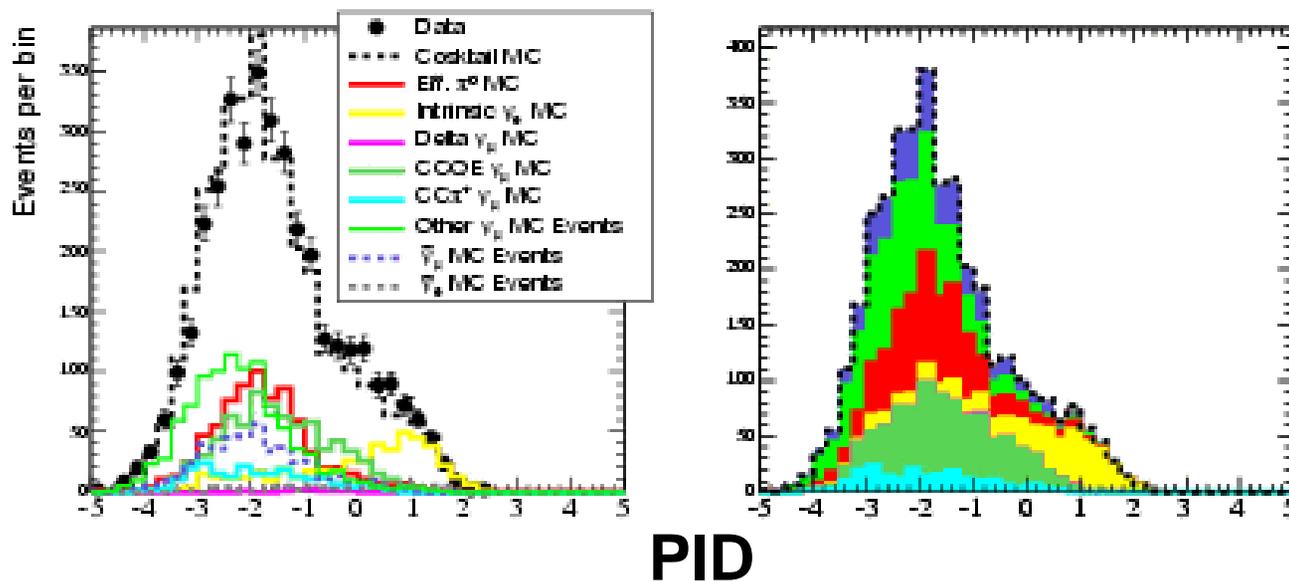
- We will use two different methods for Particle Identification: Boosted Decision Trees (**PID**) & Likelihood Cuts ( $L_{e\mu}, L_{e\pi}$ )
- See **NIM A543 (2005) 577 & NIM A555 (2005) 370** for discussions of Boosted Decision Trees, which give better performance than ANNs



# PID Distribution for Contained NuMI 1 Sub-Events

Good agreement between data and MC!

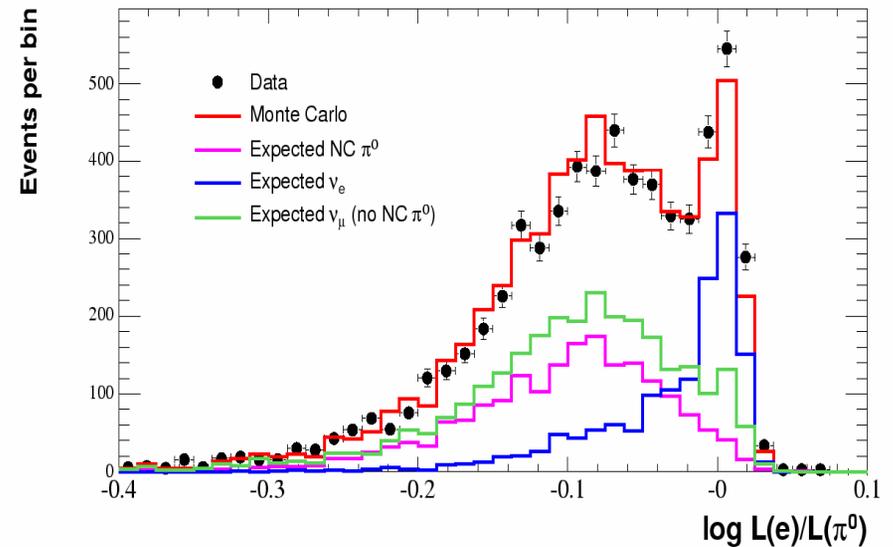
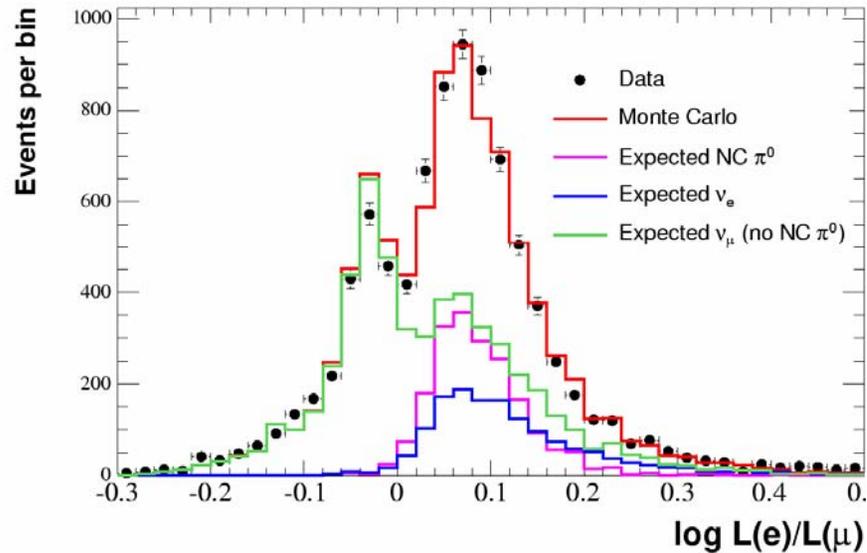
Preliminary!



# Likelihoods for Contained NuMI 1 Sub-Events

Good agreement between data and MC!

Preliminary!



# Backgrounds

- **Non-  $\nu_e$  Backgrounds**

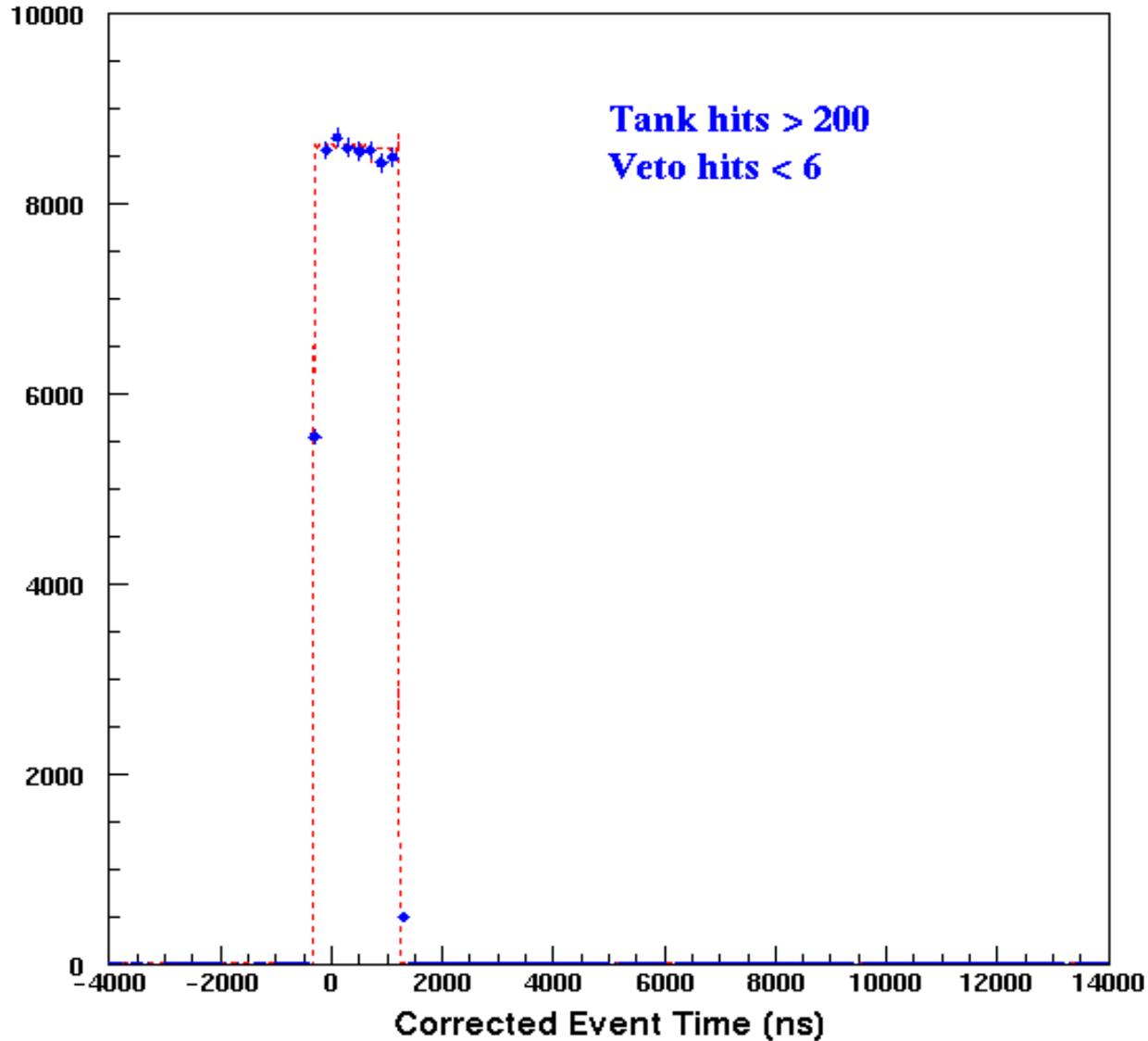
- Beam Off
- CC Inclusive
- NC  $\pi^0$
- NC  $\Delta \rightarrow N\gamma$
- NC Coherent  $\gamma$
- NC Radiative  $\gamma$

- **Intrinsic-  $\nu_e$  Backgrounds**

- From  $\mu \rightarrow \nu_e$  decay
- From  $K \rightarrow \nu_e$  decay
- From  $\pi \rightarrow \nu_e$  decay

# Beam-Off Background is Negligible!

Neutrino Signal to Cosmic-Ray Background ~ 5000 to 1!



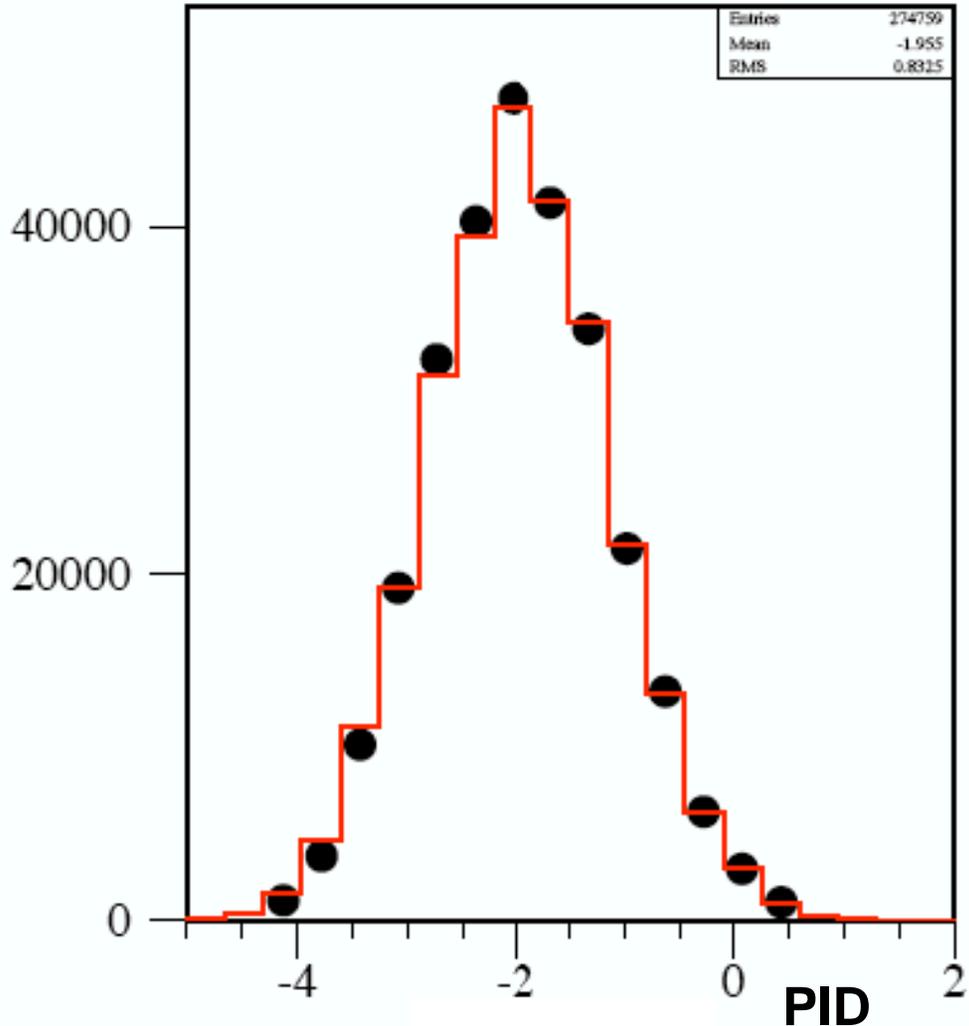
# Good PID Agreement for CC Inclusive Events

Preliminary!

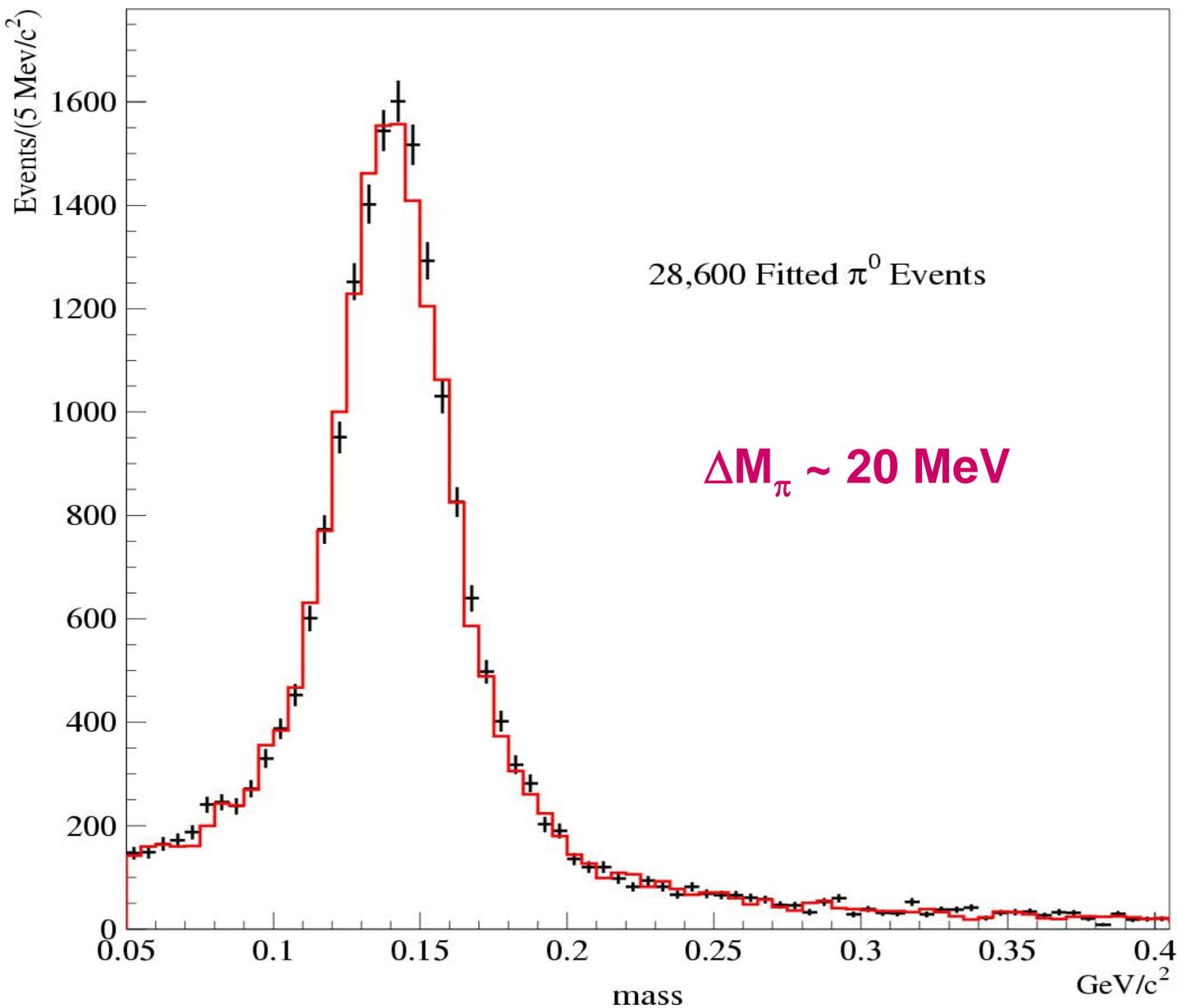
CC inclusive events  
tagged by Michel  
electrons

92% of  $\mu^-$  decay  
8% of  $\mu^-$  capture

MC is correctly  
estimating CC  
inclusive background



# MiniBooNE NC $\pi^0$ Events are Measured



# Gamma Backgrounds

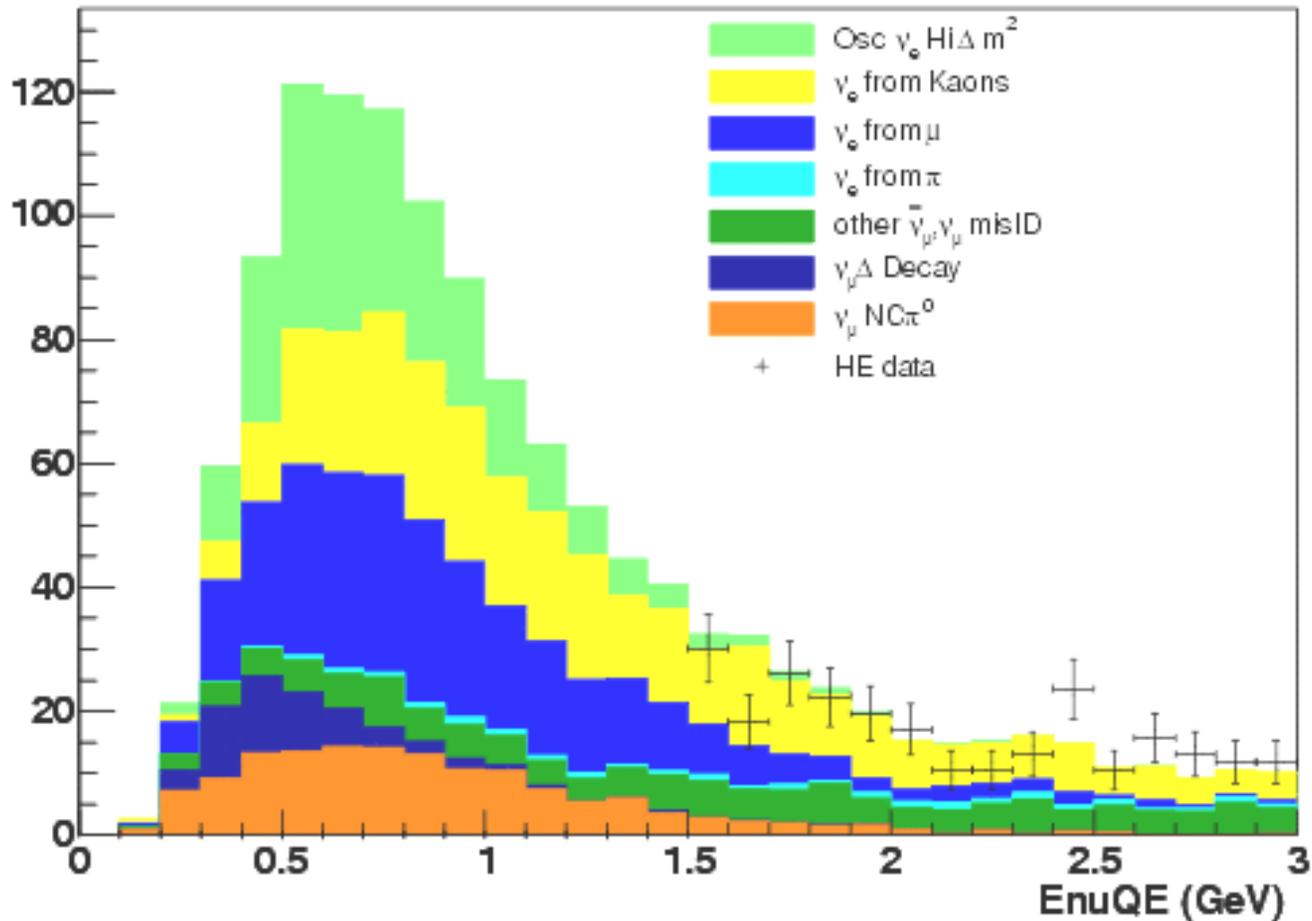
- **NC  $\Delta \rightarrow N\gamma$**  : BR  $\sim 0.56\%$  at peak; small but non-negligible background; now properly handled by modified version of NUANCE
- **NC Coherent  $\gamma$**  : Calculated from Rein & Sehgal (Phys. Lett. 104B (1981) 394) & estimated to be negligible
- **NC Radiative  $\gamma$**  : Two sources have been considered, and both are negligible: (1) Bremsstrahlung Term, where  $\gamma$  emanates from recoil proton; (2) Contact Term, where  $\gamma$  emanates from the interaction vertex

# Intrinsic $\nu_e$ Backgrounds

- **From  $\mu \rightarrow \nu_e$  decay** : Well determined from  $\nu_\mu$  CCQE events
- **From  $K \rightarrow \nu_e$  decay** : Measure kaon-induced neutrino flux at high energies & extrapolate to low energies
- **From  $\pi \rightarrow \nu_e$  decay** : Well determined from  $\nu_\mu$  CCQE events

# Expectations for 5E20 Protons on Target

Preliminary!





## Future: BooNE

If MiniBooNE confirms LSND, what new physics is there beyond the Standard Model?

With two identical detectors at different distances, many of the systematic errors cancel.

Search for sterile neutrinos by looking for oscillations of **NC  $\pi^0$**  & **NC elastic** scattering events.

# Conclusions

- MiniBooNE will soon test the LSND signal
- If the LSND signal is confirmed, then **BooNE** (with two detectors) would provide a great opportunity for Fermilab & particle physics
- Make precision measurements of the oscillation parameters
- Explore physics beyond the Standard Model (e.g. **sterile neutrinos**)