

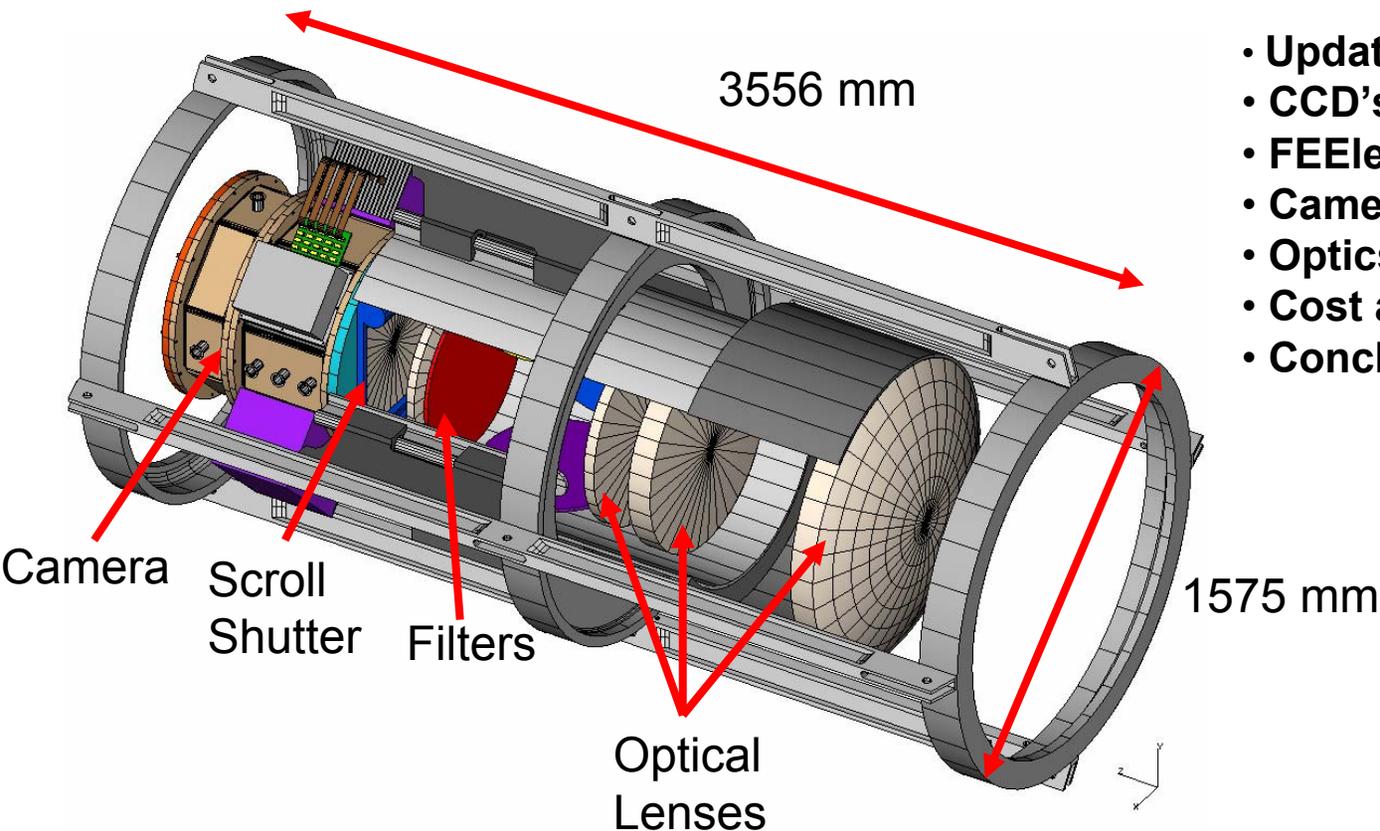


DES Technical Progress, Cost and Schedule

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SURVEY

Outline

- Update since last PAC
- CCD's
- FEElec & DAQ
- Camera Vessel/Cooling
- Optics
- Cost and Schedule update
- Conclusion





Updates

- July 2004 DES received Stage 1 approval from the Directorate, and submitted the “DES Proposal” (200 pages) to NOAO
- Juan Estrada – new Wilson Fellow joined the DES project in Sept. and is taking a leading role in the CCD testing
- Trip to CTIO Oct. 2-8: engineers figured out how to integrate DES with existing facilities, established contacts and working relationships
- Very recent Updates
 - Fermilab and LBNL have signed an MOU to initiate the CCD preproduction run
 - UC and UIUC are in the process of signing MOUs which contribute funds for DES M&S costs – these will be used mostly this year to jump start the CCD packaging development, the optics, and setting up a CCD testing center.



CCDs

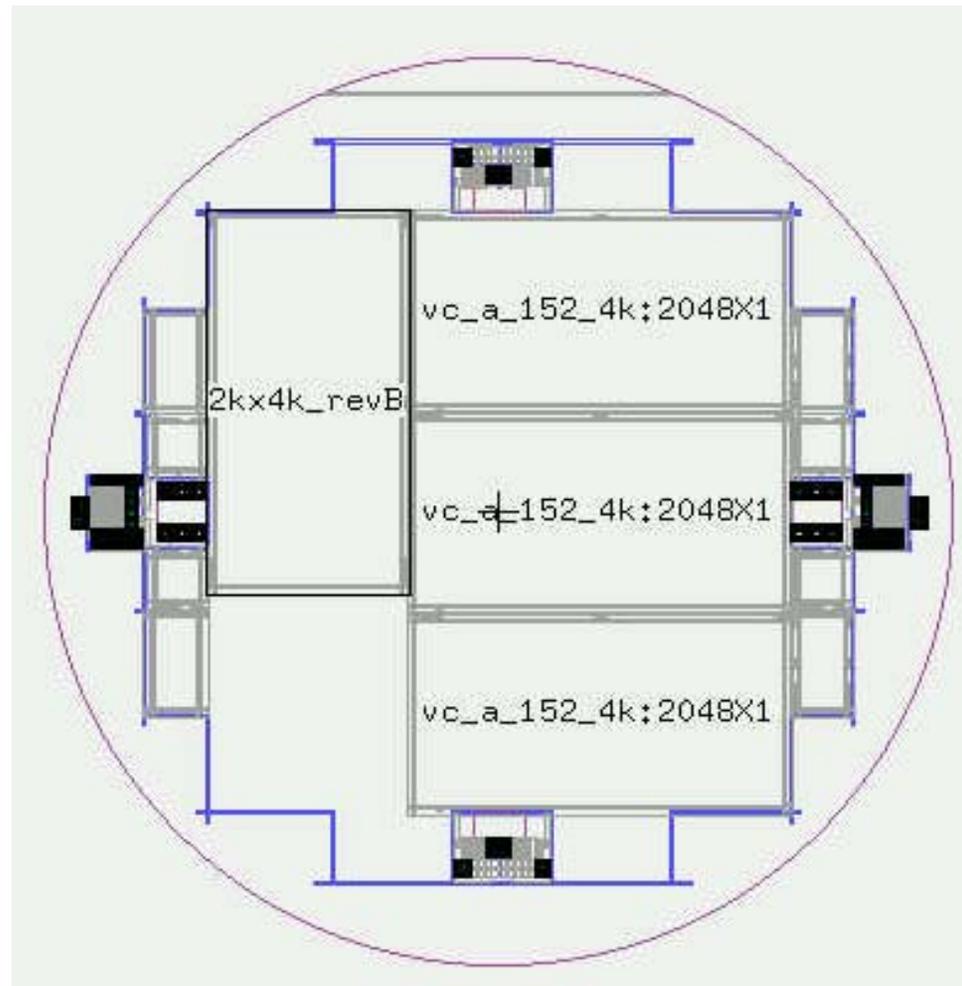
- CCD Specification Review Oct. 15th
 - Goal: generate list of specifications for the preproduction CCD masks
 - had talks on Science req., packaging, control of QE, and technical aspects of the LBNL CCDs
 - generated a spread sheet of specifications
 - final details being discussed now
- CCD Preproduction Procurement Details
 - Once we iron out the details, LBNL will design the masks (~6 wks)
 - Before submission to Dalsa we will have a final review of the masks
 - LBNL will order the wafers from Dalsa (7-9 months ARO)
 - Before processing these wafers at LBNL we come back to PPD to secure the funding (24k\$/wafer)



CCD Mask Layout

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- Each mask will contain
 - 4 2kx4k devices. These are the Image CCDs
 - 1 ~2kx2k device
 - 2 ~1kx1k devices
- For Guide and Focus chips we have options
 - 2k devices are easier to handle and package than 1k so we will fill the space with a 2k x 2k device
 - guiders need fast readout (~ 1sec) we can use sub arrays
- Small chips ~ 1k x 1k will be used as test structures in picture-frame packages for tuning up the test stands.





CCD Procurement Plans

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- Nov 2004 to Nov 2005: Phase A - Startup
 - Learn to operate CCDs,
 - begin packaging and work with R. Stover (LICK)
- Nov. 2004 to Dec. 2005: Phase B - Preproduction
 - new mask with four 2kx4k devices/wafer
 - 24 wafer run at Dalsa
 - process 2-5 wafers at LBNL
 - package and test and FNAL
 - Dec. 2005 Review Phase B CCD production and packaging results
- Dec. 2005 – July 2006: Phase C
 - process remaining Phase B wafers at LBNL
 - package and test at FNAL



CCD Procurement Plans

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- Dec. 2005 – Dec. 06: Phase D1 – Part of CCD R&D program
 - Order 3 lots of 24 wafers each, assume 20 wafers/lot survive
 - processing at Dalsa ~ 6 months
 - process and package enough wafers to determine yield (~5 wafer/lot)
 - Dec. 06 Review Phase D1 CCD production and packaging yield results and determine if contingency run is needed
- Dec. 06 – Sept. 07 Phase D2 = CCD production
 - Process Phase D wafers at LBNL as needed depending on yield
 - LBNL produces 20 tested CCDs/month for 9 months (starting 3m aro)
 - FNAL packaging and testing designed to keep up with LBNL delivery
- Dec. 06 Phase E – Contingency Run
 - if Phase D1 indicates lower than 25% yield, place order for additional wafers

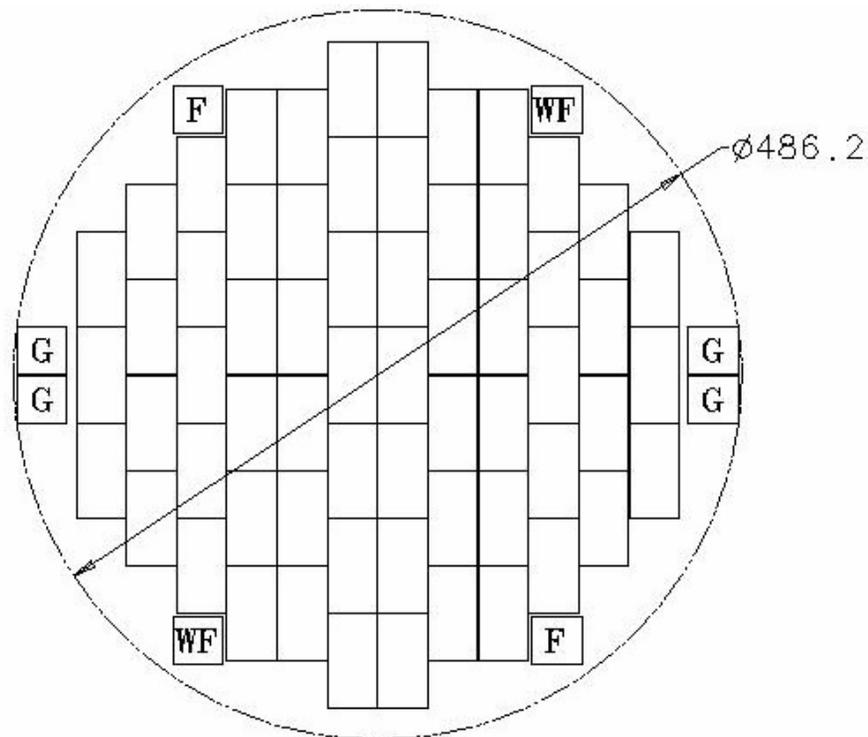
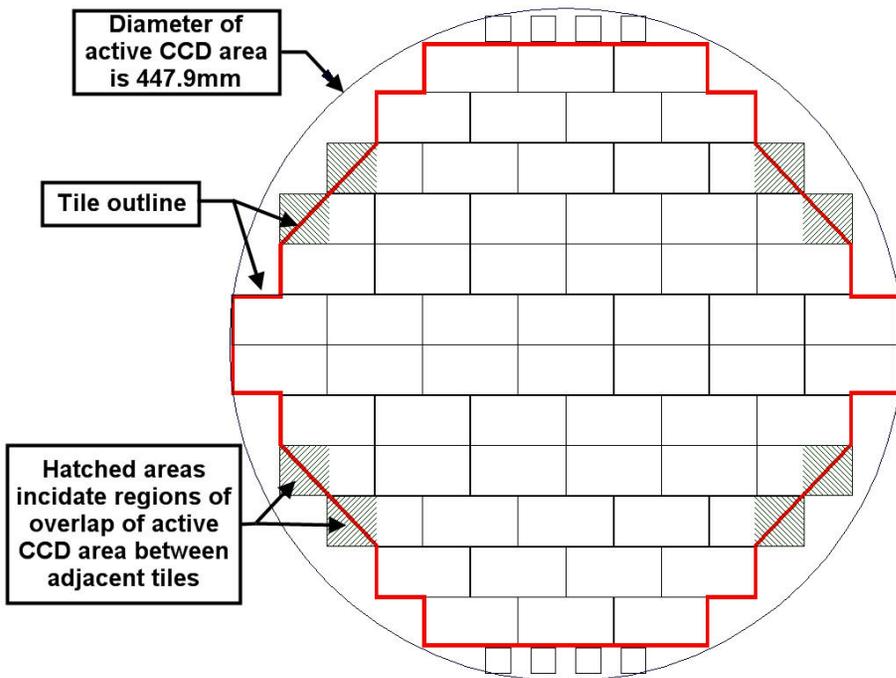


Focal plane layouts

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August 04

Nov. 04



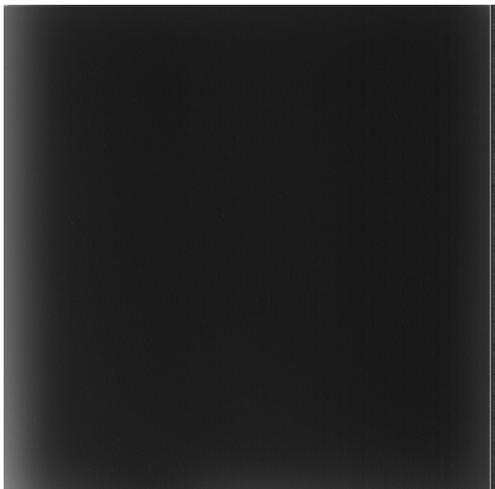
Rotated so tiles have constant declination, WF and focus sensors included, guiders slightly offset to allow for baffles



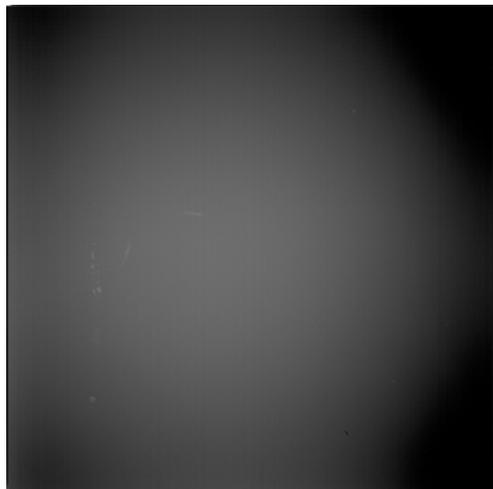
CCD testing

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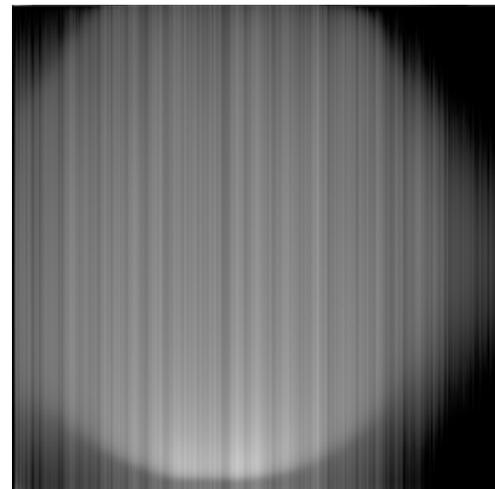
- In Sept. we borrowed a CCD and a CCD readout system from Tom Droege.
- Set up on 14th floor and then in Lab 6. It works!!
- LAB6 has most of the equipment need to test CCDs. We can borrow the equipment and dark room for a few months.



dark 10 sec



Light 10 sec



Light 100 sec, Saturated

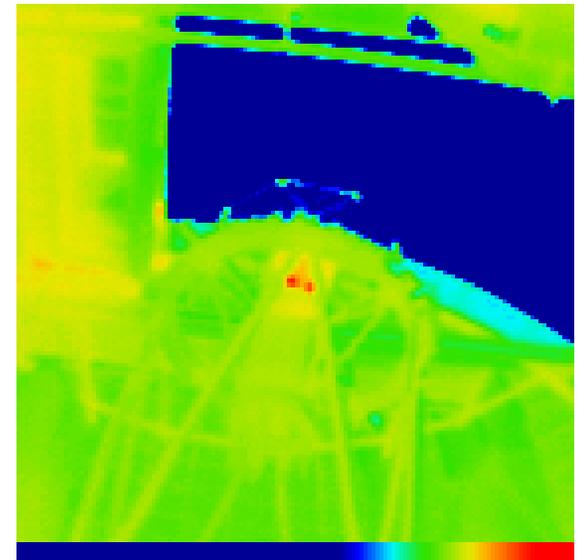


Front End Electronics and DAQ

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- NOAO is developing a Front End and DAQ system called MONSOON.
 - Front End Electronics sit in crates in PF cage
 - fibers run down the telescope trusses to the Pixel Analysis Node PC (PAN) - assembles the images
- We have two options under consideration:
 - Reference Design (in NOAO proposal): Modified/repackaged Monsoon Front End plus Monsoon software
 - Backup Solution: Monsoon Front End Electronics (DHE) + Monsoon software
 - this is what we will use for testing CCDs in the near future and probably for at least the next year (or two)
- **Issues for Prime focus cage are heat and space. Cost and schedule are also considerations**

Thermal image

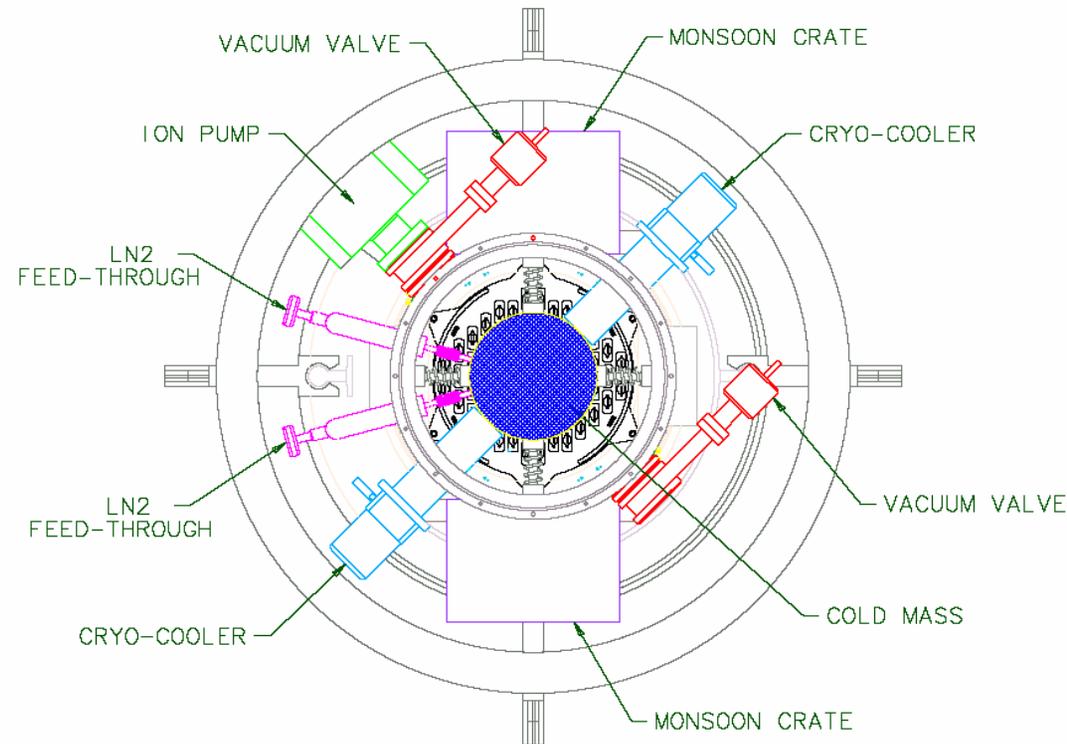




Cooling System

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- Cooling with Cryo-Coolers
 - Design includes two
 - Both operate during cool-down
 - Require one in normal operations
 - Vibration analysis is ongoing
 - Use of Vibration Dampener is planned
- Cooling Reserve Options
 - LN2 Reservoir
 - 3-4 hour reserve is possible
 - Implementation is more complex than a cold mass
 - Copper Cold Mass
 - 30 minute reserve ($dT=20$ K)
 - Easier to maintain a consistent temperature distribution



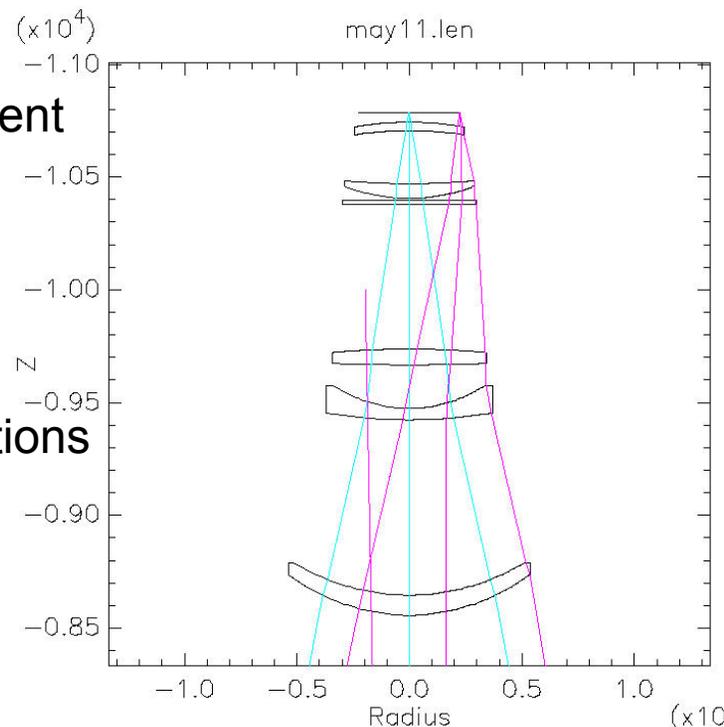


Corrector Optics

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Chicago and hopefully UCL will take the lead on the Optical design

- Steve Kent and Mike Gladders developed our current optical design (May 04)
- UCL has an Optical Science Center & experience procuring corrector optics.
- UCL reviewed our design and presented some options for future optimization
 - reverse the aspheric surface (manufacturing)
 - remove one element, have a flat window
- UCL also obtained independent cost estimates.
 - Glass 600k\$, 5 months – agrees with our est.
 - Figuring 600k\$, 1 year delivery (400k\$ and 6m less than in our schedule)

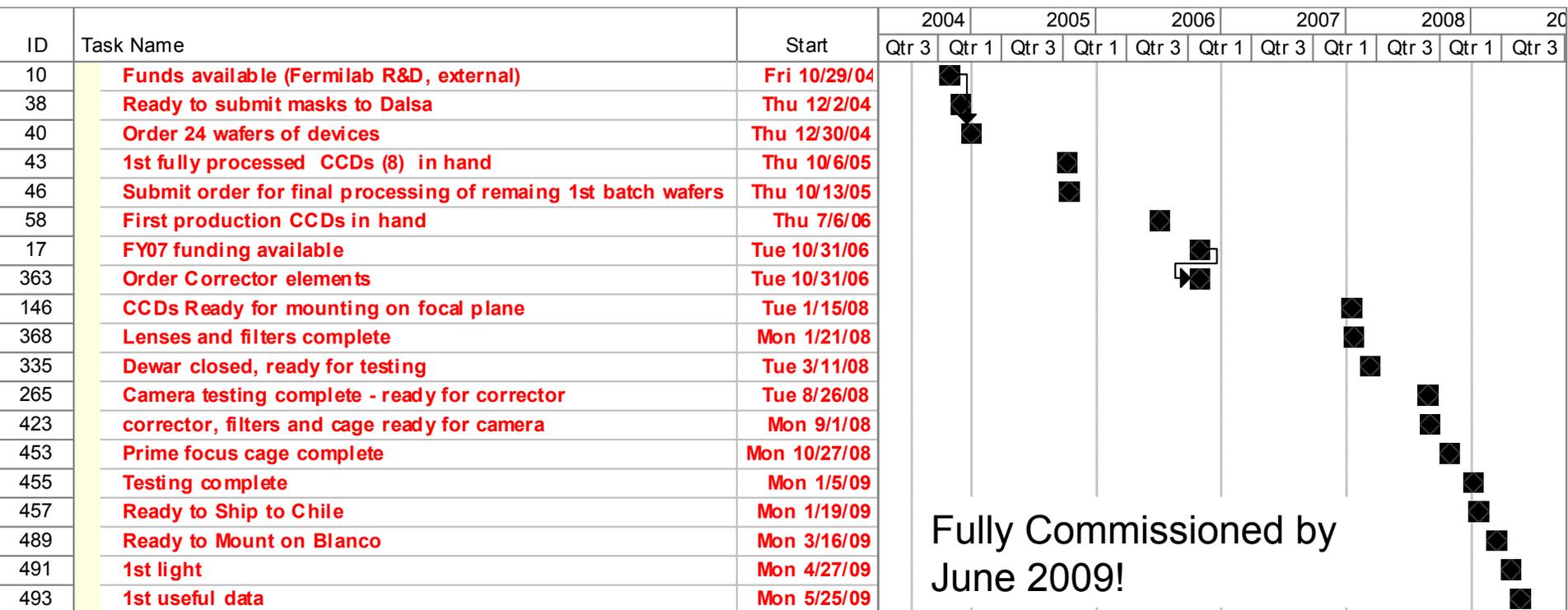




Schedule Milestones

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- Optics and CCDs are the most Challenging tasks
- Optics are critical path: Order glass in FY06, Figuring/polishing in FY07





Total Cost profile in Then Yr \$

(excluding institutional overhead)

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This corresponds to the profile discussed in the responses to the PAC questions for the June 04 meeting.

The optics procurements are funding limited and set the critical path.

	FY04	FY05	FY06	FY07	FY08	FY09	Total
M&S	0	1,097	3,321	3,362	520	68	8,368
M&S Contingency	0	0	379	338	1,580	782	3,079
Total M&S	0	1,097	3,700	3,700	2,100	850	11,447
Labor	609	1062	955	1150	560	299	4,635
Labor Contingency	304	531	478	575	280	150	2,318
Total Labor	913	1593	1433	1724	839	449	6,953
Total (M&S + Labor)	913	2,690	5,133	5,424	2,939	1,299	18,400

Estimated overhead if all funds spent at FNAL is 4.1M\$ in then year \$

This profile has us ready for observations by June 2009.



Plans for FY05

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- Initiate the Preproduction CCD order – Done
- Learn to test CCDs – in progress
 - Testing started (all borrowed equipment)!
 - Fabricate our own test Dewar – drawings out for bids
 - Fabricate “personality board” for testing LBNL CCDs with TD readout system – in layout
- Learn to package CCDs – waiting for MOUs to be signed
 - order a small number of devices ~ off the shelf at LBNL (3m ARO)
 - begin procuring the CCD packaging hardware
- Hope to have a 2k x 4k CCD (phase A) packaged and tested at FNAL by early summer 2005
- Initiate Prototype efforts for FE electronics, filter and shutter
- Optical design trade studies and firmer quotations

- We will rely on UIUC, UC and Fermilab for the R&D funds



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Conclusions

- Money starting to be available, project is getting underway!



Dark Energy Survey (E939)

Funding status and reviews

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- DES Status as of July 2004
- NOAO response to the DES proposal
- Status of financial commitments to the DES
- Securing the remaining funds
- Dark Energy Task Force and fund raising



DES Status as of July 2004

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- DES had prepared a proposal that included a design and preliminary cost estimate of DECam (the survey instrument).
- It was reviewed by the Laboratory and considered by the PAC.
- June PAC meeting outcome:
 - Recognized importance of the DES science and its relevance to the Fermilab scientific program
 - The Director authorized the DES to submit the proposal to NOAO
 - The Director granted Stage I approval



DECam cost and funding plan as of July 2004

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- DECam estimated cost \$18.4 M*
 - M&S (cash) \$11.4 M
 - Staff costs (exclusive of scientists) \$7.0 M

*without indirect costs of ~\$4.1 M

- DECam Funding plan
 - Fermilab M&S request \$7.4 M
 - Fermilab staff request \$ 6.4 M
 - UIUC-HEP staff request \$ 0.6 M
 - Cash from “Other Sources” \$ 4.0 M

- “Other sources”: NSF and private sources



NOAO Review and outcome

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- The NOAO Director appointed the BIRP (Blanco Instrumentation Review Panel) to review our proposal:
 - The BIRP Recognized the importance of our science and
 - recommended that the DES collaboration form the instrumentation partnership described in the AO with NOAO.
- The Director accepted the recommendations and asked the CTIO Director and DES Project Director to draft a Memorandum of Agreement for signature by the NOAO and Fermilab Directors. It will provide:
 - DECam, the software and archive to NOAO
 - the observing time to the DES Collaboration
 - the telescope improvements and infrastructure for DECam



Financial Commitments to the DES

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SURVEY

- The UIUC and Chicago MOU's are ready for signature. They provide DES with:
 - \$ 400K (all available in FY05) from UIUC
 - \$ 500 K (\$300 K available in FY05) from Chicago
 - The indirect cost as an FNAL contribution
- Fermilab has given DECam an FY05 budget that :
 - provides funding for the masks and the wafers and enables the Fermilab-LBNL MOU for the CCDs.
 - the DECam development program can proceed with the university funds
- DES now needs to raise \$3.1 M from “other sources”



DES presentation to the NSF and outcome

DARK ENERGY
SURVEY

- We presented the DES plan to the NSF in September. It was well received, but later we learned that:
 - NSF will not provide funds for the instrument.
 - The NSF will consider a DES Data Management proposal if the Dark Energy Task Force recommends an intermediate term Dark Energy program that contains a survey like the DES.
- The DES elected to delay submitting their Data Management proposal to the NSF until after the DETF issues its preliminary report (early summer?).



DES presentation to DOE and outcome

DARK ENERGY
SURVEY

- We presented the DES plan to DOE/HEP in September and it was well received. We learned that:
 - DOE/HEP will allow Fermilab to spend R&D funds in FY05 and FY06 for DES development.
 - DOE/HEP will use the DETF report to request CD-0 from DOE for the intermediate term Dark Energy Program .
 - FY07 is the earliest budget year for the Dark Energy Program
 - The DES project could be base-lined (CD-1) in late summer 05



“Other Sources”: new partners

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- University College London, Cambridge and Edinburgh have proposed to join the DES. Lahav, Sutherland, Peacock and Doel submitted a Statement of Interest to PPARC. It requests 1.2 M pounds.
 - They propose to provide the optical elements as an in kind contribution. This will reduce our cash needs by > \$ 1 M.
 - They are interested in weak and strong lensing, photo-z’s, and galaxy evolution. They strengthen our science effort.
 - UCL is already working on the DECam optics design and photo-z’s on a provisional basis.
 - They expect to hear on whether they will be invited to submit a proposal to PPARC on November 18.



Other possibilities for new partners

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SURVEY

- We are searching for new partners that can afford the ~\$500 K admission price. Each new partner must reduce our need for cash by ~ \$500 K.
- We have contacted several interested US Universities that understand the admission fee.
- We have had more indefinite contacts as well
- We will also search for private donors.
- If the UK consortium proposal is approved we will only need \$2 M in cash, less any reduction in contingency that our cost studies justify.



Summary of the DES financial status

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- DES looks forward to the Dark Energy Task Force report and expects that it will get a place in the intermediate term Dark Energy program.
- If the DES gets a place then a joint DOE/NSF task force will review the DES. This step will be competitive with other intermediate term projects that propose the same goals. We are confident that we can succeed.
- We are confident that we can raise all of the “Other Funds”. We will still have to pass DOE reviews for CD-1, CD-2, and CD-3 and the appropriate NSF reviews.
- We should know our fate by late summer.



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SURVEY

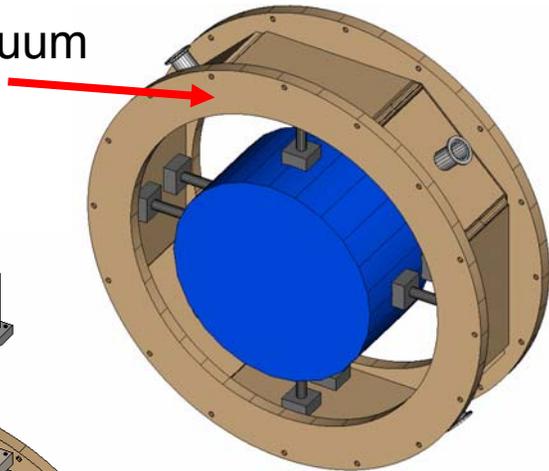
Extra slides



Camera Vessel

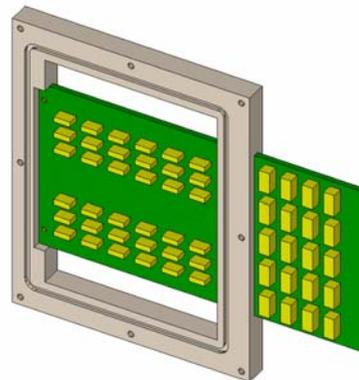
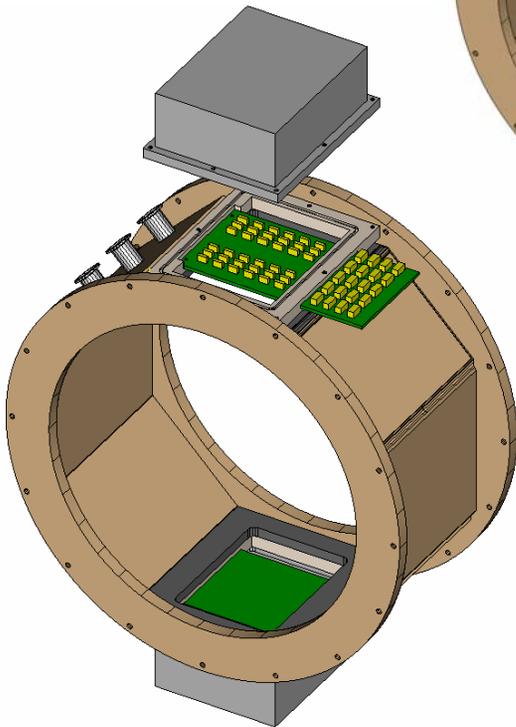
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SURVEY

Cooling/ Vacuum
spool piece



Camera is separated into two spool pieces: one for signal feed throughs one for cooling and vacuum services

Removal of cooling spool piece allows access to back of focal plane and cables



- Vacuum feed through board brings signals out of cryostat



Cost Estimate breakdown

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		M&S (\$K)	Labor (\$K)	Total (\$K)
1	Dark Energy Survey Instrument	8,368	4,635	13,003
1.1	Management	235	189	423
1.2	Dark Energy Instrument Construction	7,734	3,692	11,426
1.2.1	CCDs	2,155	25	2,180
1.2.2	CCD Packaging	287	607	895
1.2.3	Front End Electronics	595	640	1,235
1.2.4	CCD Testing	579	364	942
1.2.5	Data Acquisition	167	299	466
1.2.6	Focal Plane, Camera Vessel	163	366	529
1.2.7	Cooling	427	160	587
1.2.8	Corrector	2,463	121	2,584
1.2.9	Primary Cage mechanics, corrector housing	618	779	1,397
1.2.10	Auxiliary Components	146	164	309
1.2.11	Final Assembly and Testing	135	167	302
1.3	Survey Strategy	399	539	938
1.4	Telescope Improvements	0	0	0
1.5	Commissioning (at CTIO)	0	134	134
1.6	System Integration	0	82	82
1.7	Dark Energy Science	0	0	0
	Contingency	3,080	2,318	5,397
	Total cost (w/o overhead)	11,447	6,953	18,400

Costs in
Then Year \$
w/o overhead

\$22.5M
inc. overhead



CCD Testing Setup at LAB 6

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